

# Coronal loop seismology using the Markov Chain Monte-Carlo techniques

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# Outline

- 1 Bayesian approach
  - Uncertainties estimation
  - Model comparison
  - Posterior predictive check and forecasting
- 2 Markov Chain Monte Carlo
- 3 Application to the coronal loop seismology
- 4 Summary

# The Bayes theorem

## Improving knowledge

**Knowledge before measurements** [Prior PDF -  $P(\theta)$ ]  $\otimes$

**Measurements** [The Likelihood -  $P(D|\theta)$ ]  $\Rightarrow$

**Improved knowledge** [Posterior PDF -  $P(\theta|D)$ ]

# The Bayes theorem

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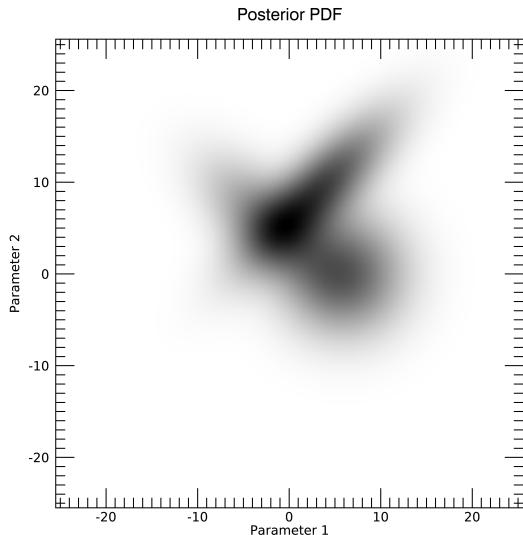
## Evidence

$$P(D) = \int P(D|\theta)P(\theta)d\theta$$

**Evidence** is the quantitative measure of how good the model  $M$  describes observational data  $D$ .

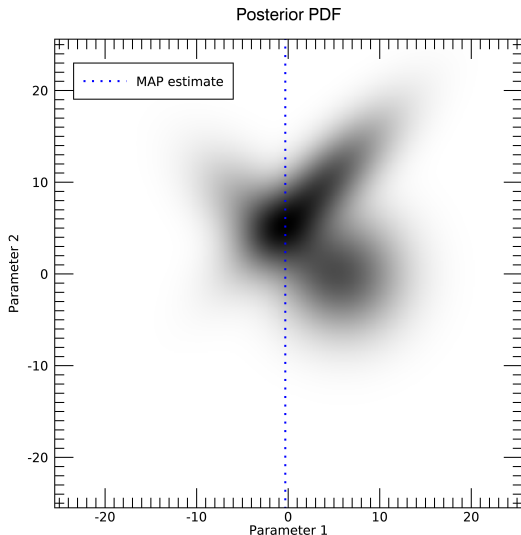
# Extracting information from the Posterior PDF

## Posterior PDF



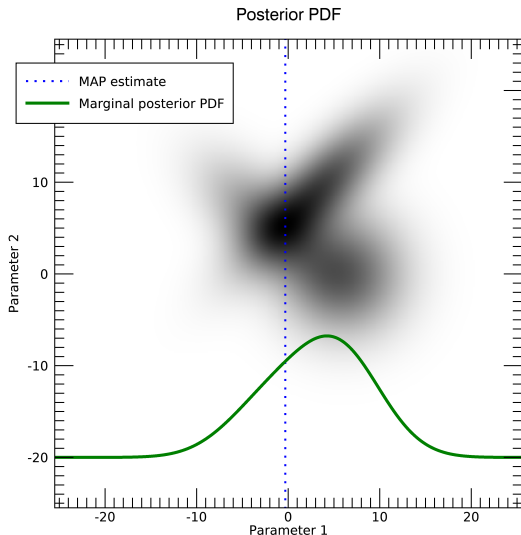
# Extracting information from the Posterior PDF

## Maximum A Posteriori (MAP) estimate



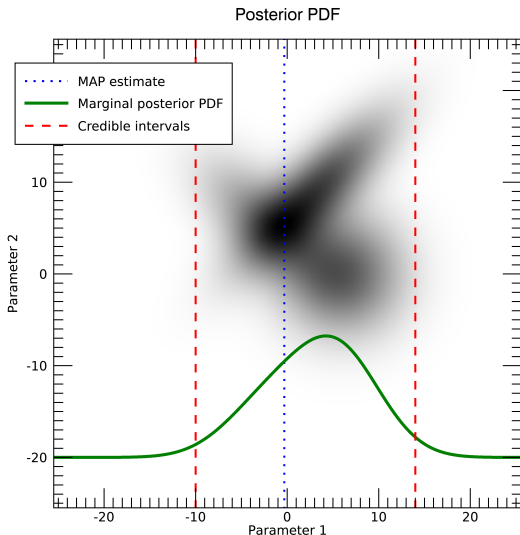
# Extracting information from the Posterior PDF

## Marginal Posterior PDF



# Extracting information from the Posterior PDF

## Credible intervals



# Model comparisons

Bayes factor

## Evidence

$$P(D) = \int P(D|\theta)P(\theta)d\theta$$

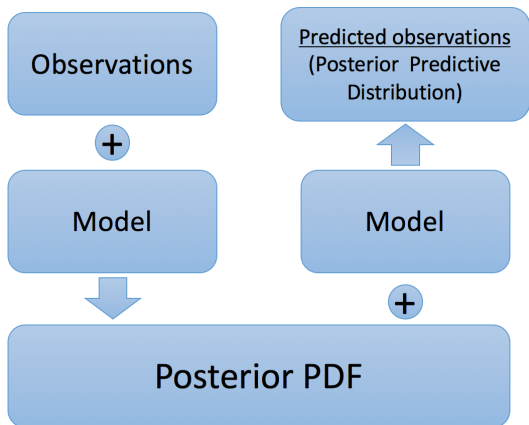
**Evidence** is the quantitative measure of how good the model  $M$  describes observational data  $D$ .

$$B_{12} = \frac{\int P(D|\theta, M_1)P(\theta|M_1)d\theta}{\int P(D|\theta, M_2)P(\theta|M_2)d\theta}$$

$B_{12}$	$2 \ln B_{12}$	Model 1 evidence
1 - 3	0-2	not worth more than a bare mention
3 - 20	2-6	Positive
20 - 150	6-10	Strong
>150	>10	Very strong

# Posterior Predictive Distribution

## Predictive Check and Forecasting



# The curse of dimensions

The determination of the credible intervals requires the computation of the marginal PDFs:

$$P(\theta_i|D) = \int P(\theta_1, \theta_2, \dots, \theta_N|D) d\theta_{k \neq i}$$

## The curse of dimensions

- Analytical integration is often impossible;
- Direct numerical integration every additional parameter increases the computation time by several order of magnitudes.

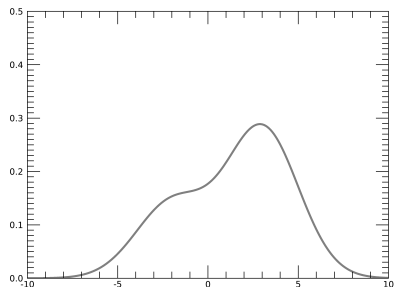
## Solution

- Sample the posterior distribution;
- Approximate marginal PDFs by histograms.



# Markov Chain Monte-Carlo

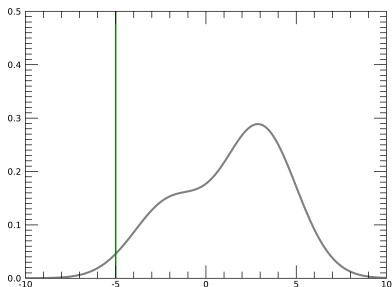
How does MCMC sampler work?



# Markov Chain Monte-Carlo

How does MCMC sampler work?

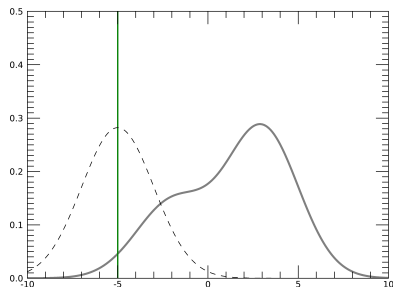
- 1 Set a walker in the parameter space



# Markov Chain Monte-Carlo

How does MCMC sampler work?

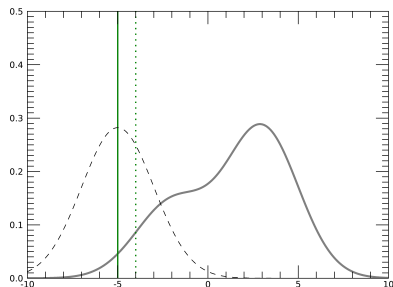
- 1 Set a walker in the parameter space
- 2 Simulate a new position from the **proposal distribution**



# Markov Chain Monte-Carlo

How does MCMC sampler work?

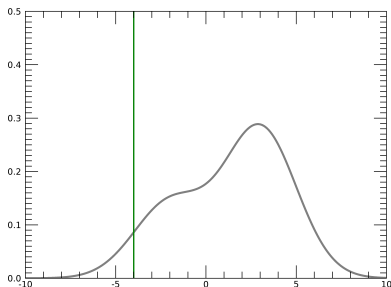
- 1 Set a walker in the parameter space
- 2 Simulate a new position from the **proposal distribution**
- 3 Accept or reject the step according a **special rule**



# Markov Chain Monte-Carlo

How does MCMC sampler work?

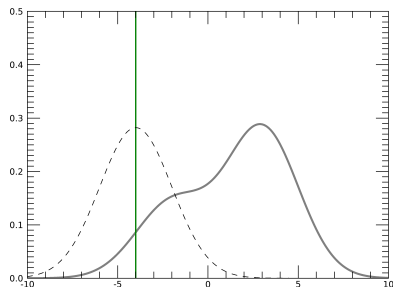
- 1 Set a walker in the parameter space
- 2 Simulate a new position from the **proposal distribution**
- 3 Accept or reject the step according a **special rule**
- 4 Return the current position as a sample from **the target distribution**



# Markov Chain Monte-Carlo

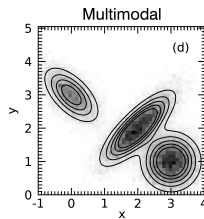
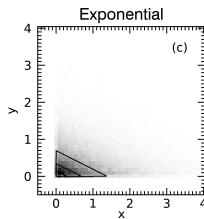
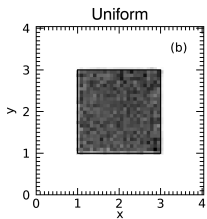
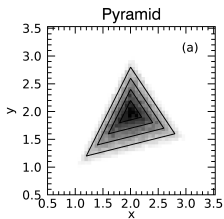
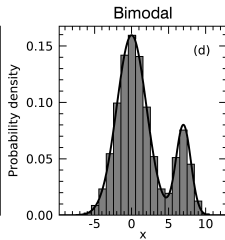
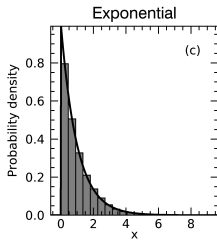
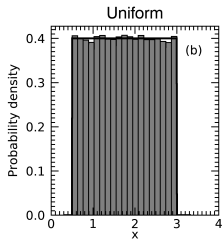
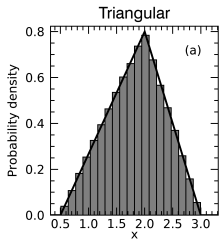
How does MCMC sampler work?

- 1 Set a walker in the parameter space
- 2 Simulate a new position from the **proposal distribution**
- 3 Accept or reject the step according a **special rule**
- 4 Return the current position as a sample from **the target distribution**
- 5 Go to 2



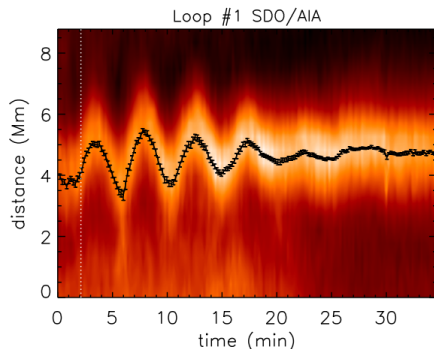
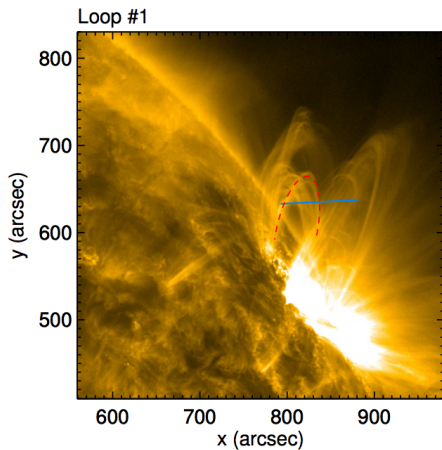
# MCMC

## Tests



# Observations

7 January 2013 06:38:11 UT

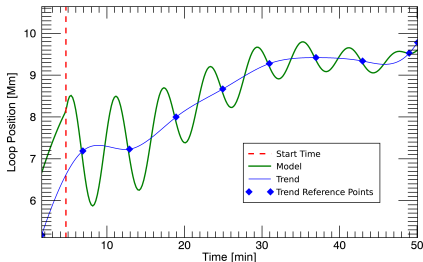




# Models of decaying kink oscillations

## Parameters:

- $A_0$  – amplitude
- $P$  – period
- $\tau$  – decay time
- $t_0$  – starting time
- $x_0$  – initial displacement



$$y(t) = \text{Trend}(t) + \begin{cases} A_0 e^{-\left(\frac{t-t_0}{\tau}\right)^n} \sin \left[ \frac{2\pi(t-t_0)}{P} + \arcsin\left(\frac{x_0}{A_0}\right) \right], & t \geq t_0 \\ x_0, & t < t_0 \end{cases}$$

## Models

- Exponential decay model ( $M_e$ ):  $n = 1$
- Gaussian decay model ( $M_G$ ):  $n = 2$

# The model

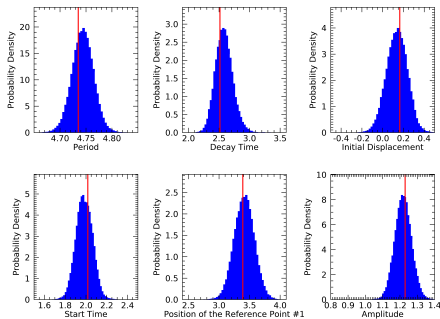
- Observables  $D = [Y_1, Y_2, \dots, Y_{N_d}]$
- Unknown model parameters  $\theta = [A_0, P, \tau, \dots]$
- The model  $Y_i^m(\theta) = y(\theta) + N_i(0, \sigma^2), 1 < i < N_d,$ 
  - ▶  $N_d$  — number of data points
  - ▶  $N(0, \sigma)$  — white noise with zero mean and dispersion of  $\sigma^2$

## The likelihood

$$P(D|\theta) = \frac{1}{(2\pi\sigma^2)^{\frac{N_d}{2}}} \prod_{i=1}^{N_d} \exp \left\{ -\frac{[Y_i - Y_i^m(\theta)]^2}{2\sigma^2} \right\}$$

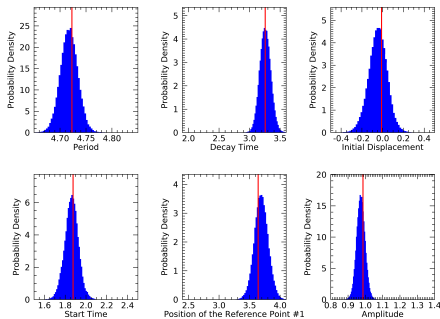
# Parameters Inference

## Exponential decay



$$P(D|M_e) = 1.5 \times 10^{28}$$

## Gaussian decay

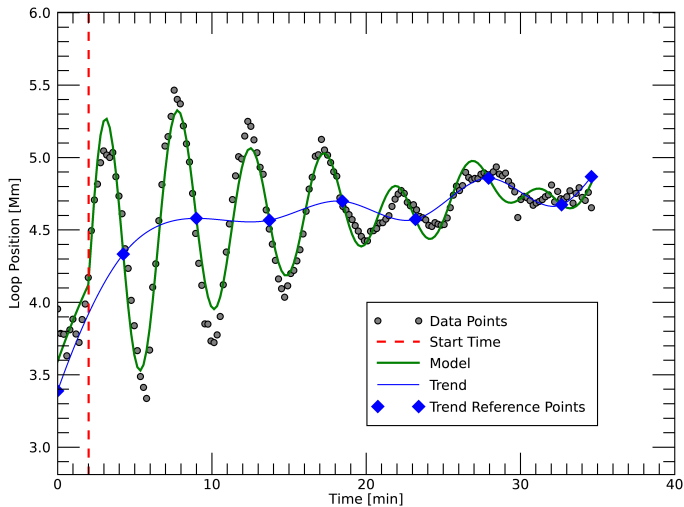


$$P(D|M_G) = 1.0 \times 10^{47}$$

$2 \ln B_{e,G} = 86.7$  (Very Strong evidence for Gaussian decay)

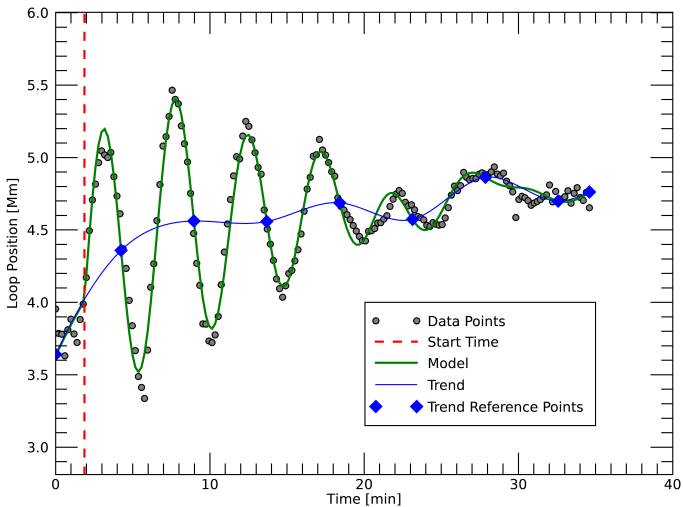
# Fitting results

## Exponential decay



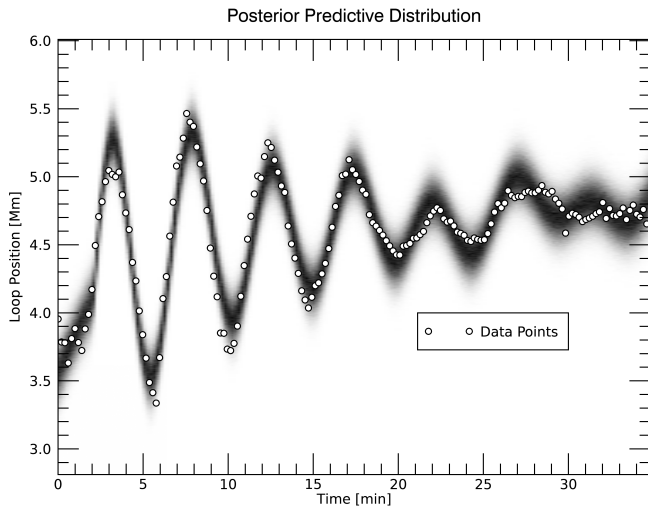
# Fitting results

## Gaussian decay



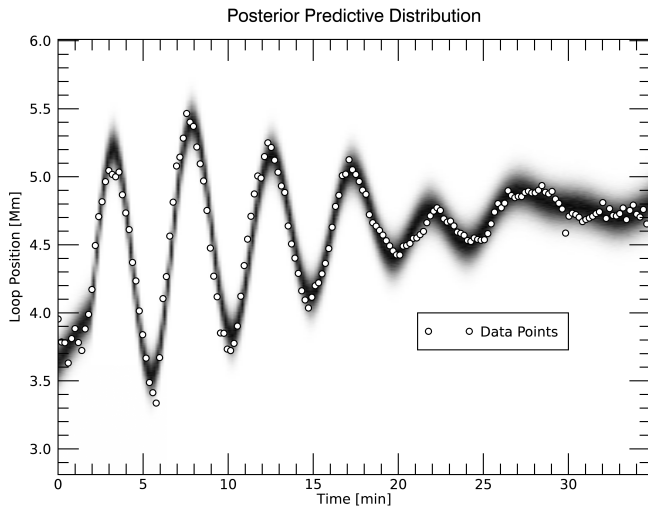
# Posterior Predictive Distribution

Exponential decay



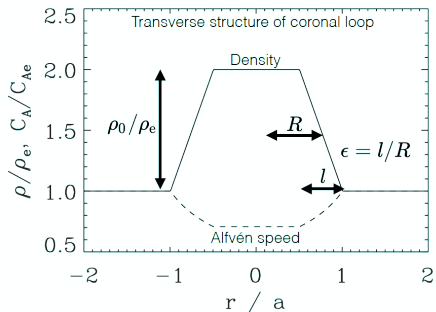
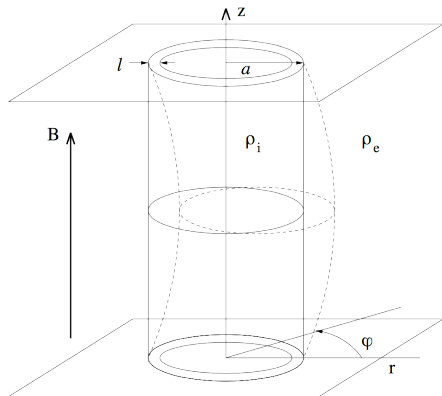
# Posterior Predictive Distribution

Gaussian decay



# Straight cylinder model of a coronal loop

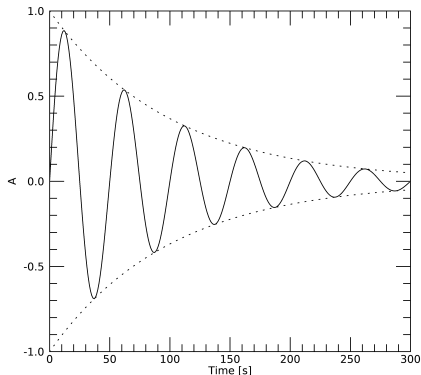
with the transition layer





# Exponential decaying profile<sup>1</sup>

$$D(t) = \exp \frac{t}{\tau}, \tau = \frac{2P}{\pi \epsilon \kappa}$$



## Unknown parameters

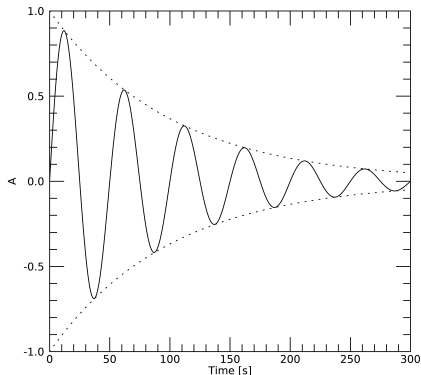
1  $\epsilon = l/R$

2  $\kappa = \frac{\rho_0 - \rho_e}{\rho_0 + \rho_e} = \frac{\rho_0/\rho_e - 1}{\rho_0/\rho_e + 1}$

<sup>1</sup>[Goossens et al., 2002, Ruderman and Roberts, 2002]

# Exponential decaying profile<sup>1</sup>

$$D(t) = \exp\left(-\frac{t}{\tau}\right), \tau = \frac{2P}{\pi\epsilon\kappa}$$



## Unknown parameters

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## Observables

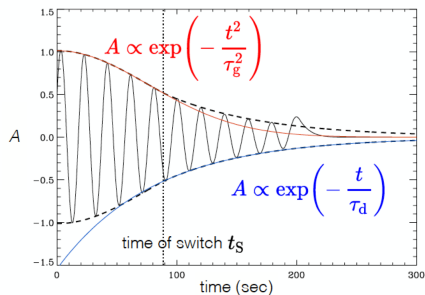
1 Decay time  $\tau$

<sup>1</sup>[Goossens et al., 2002, Ruderman and Roberts, 2002]

# General decaying profile <sup>2</sup>

$$D(t) = \begin{cases} \exp\left(-\frac{t^2}{2\tau_g^2}\right), & t \leq t_s \\ A_s \exp\left(-\frac{t-t_s}{\tau_d}\right), & t > t_s \end{cases}$$

$$\tau_g = \frac{2P}{\pi\kappa\epsilon^{1/2}}, \quad \tau_d = \frac{4P}{\pi^2\kappa\epsilon}, \quad \tau_s = \tau_g^2/\tau_d$$

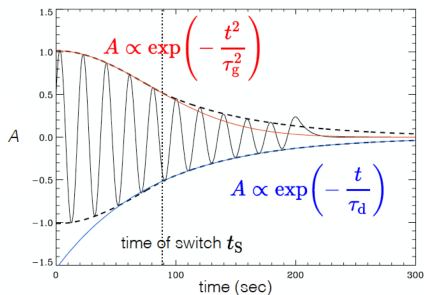


<sup>2</sup>[Hood et al., 2013, Pascoe et al., 2013]

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## Unknown parameters

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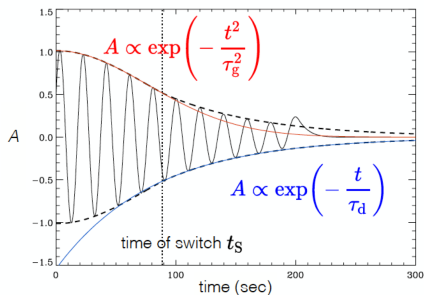
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## Unknown parameters

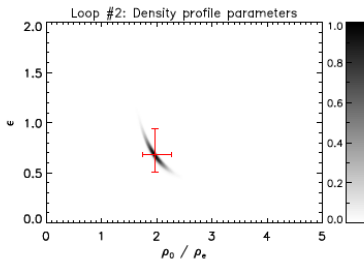
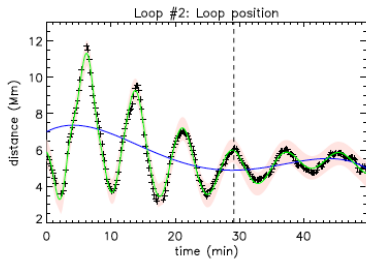
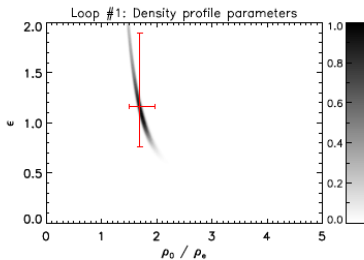
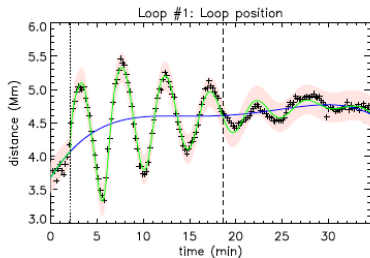
- 1  $\epsilon = l/R$
- 2  $\kappa = \frac{\rho_0 - \rho_e}{\rho_0 + \rho_e} = \frac{\rho_0/\rho_e - 1}{\rho_0/\rho_e + 1}$

## Observables

- 1 Gaussian decay time  $\tau_g$
- 2 Exponential decay time  $\tau_d$

<sup>2</sup>[Hood et al., 2013, Pascoe et al., 2013]

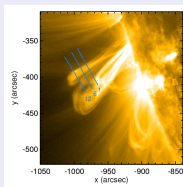
# Full Inversion<sup>3</sup>



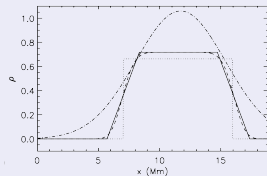
<sup>3</sup>[Pascoe et al., 2017a]

# Density profile inversion from the EUV intensity <sup>4</sup>

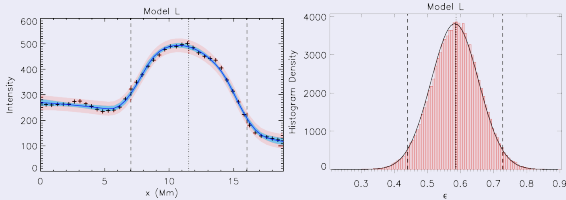
SDO/AIA 171 Å



Inferred density profiles



MCMC inference results



<sup>4</sup>[Pascoe et al., 2017b]

# Summary

MCMC is a universal method for solving inverse problems.

- Output: probability distribution function for the model parameters
  - ▶ Most credible values
  - ▶ Robust error estimations
- Comparison of the competing models using the Bayes factor
  - ▶ Transparent accounting for the number of parameters
  - ▶ Comparison of rather the models than best fits.
- In combination with advanced physical model, it allows us to extract as much reliable information as possible from the observational data
- We inferred transverse density structure of several coronal loops
  - ▶ Seismological inversion from kink oscillations
  - ▶ Inversion from EUV intensity profiles



Thank you for your attention!

# References



Goossens, M., Andries, J., and Aschwanden, M. J. (2002).

Damping of coronal loop oscillations by resonant absorption of quasi-mode kink oscillations.

In Wilson, A., editor, *Solar Variability: From Core to Outer Frontiers*, volume 506 of *ESA Special Publication*, pages 629–632.



Hood, A. W., Ruderman, M., Pascoe, D. J., De Moortel, I., Terradas, J., and Wright, A. N. (2013).

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